

Optical Crystals Demo

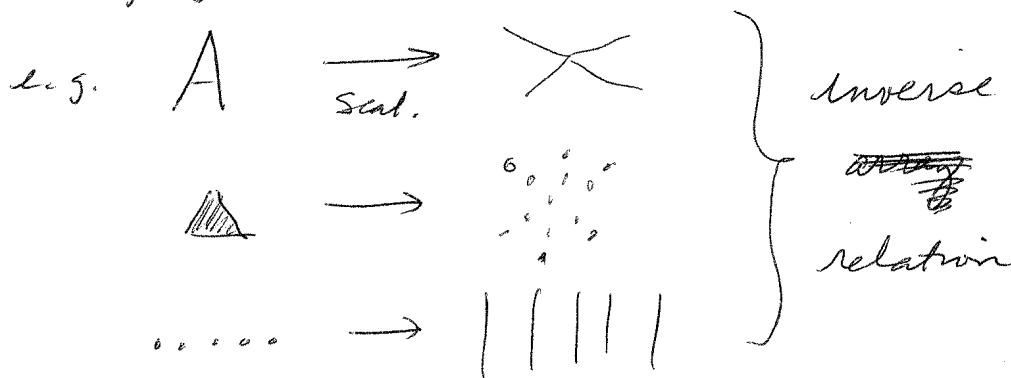
- use slides in General Microscopy
- try to guess shapes using the real optical crystals
- It's hard: from this we conclude that microscopy is superior to LS when you can get away with it.

But....

- floppy objects better in LS
- averaging better in LS
- dynamics / kinetics
- quantization
- Theory

So let's look closer at the optical crystals observations

- Square arrays resulted in square diffraction
- hex " " " hex "
- intensity of points modulated by shape of particle



But when we didn't know the answer, it was often hard. This is due to Phase Problem

Phase Problem: Cantor & Schimmel, Ch. 13

Image \Leftrightarrow Scattered Electric field (perfect)
 $| \text{Scattered E field} |^2 \Leftrightarrow$ What you see (imperfect)
 (intensity)

$$I = S \cdot P$$

↗ ↗
 form factor (e.g. \propto for A)
 structure factor
 (e.g. square pattern for square array of A's)

High Polymer Physics Prize: all these prizes show us that there is powerful motivation to "see" molecules through scattering, despite the hardships. We may expect that some mathematical apparatus will be required to figure out why $\frac{\partial}{\partial \theta} \frac{\partial}{\partial \phi} = \Delta$

Light Scattering: Outline

Observations, including diffraction which is a special case of scattering by regularly aligned objects

Small Molecules

Gas

- A) Particle picture for a single particle.
- B) Dilute picture: "matrix" type summation of electric fields
- C) Molecular weights

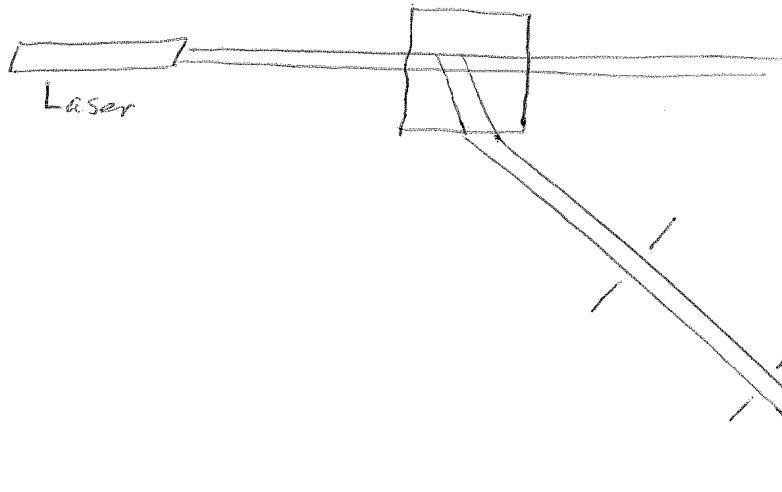
Two-component Solution (adapt from gas)

- A) Simple adaptation from gas
- B) Thermodynamic Viewpoint

Large Molecules

Preface: LS is relatively ignored by chemists who are mostly interested in another way that light interacts with matter: absorption. Nevertheless, scattering is as basic - and as important - as absorption.

The Nature of the LS experiment

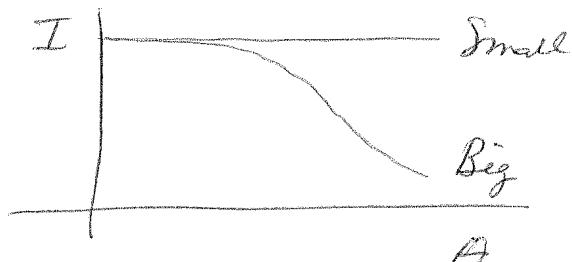


Snell's Law
 Marching Demo
 Doppler page
 LS.2±

Experiential Experimental Observations

① Other things being equal, big particles scatter more than little ones

② Angular Dependence

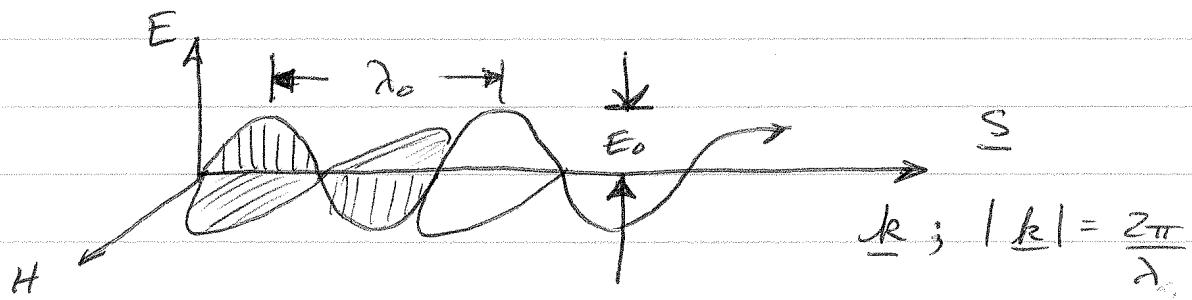


③ Blue sky ... but red sunsets
 Polarized Sunlight
 Sky seems far away ... but This is frcky!

④ All that optical crystal stuff

A word first about light.

We already discussed that light produces rapidly (10^{14} - 10^{15} Hz) oscillating electrical fields. But there is more to it than that.



Poynting vector

$$\underline{S} = \frac{c}{4\pi} \underline{E} \times \underline{H}$$

\underline{S} describes the instantaneous rate of energy flow at a point

Average over time to give I as energy per sec.
per squared centimeter in direction of propagation

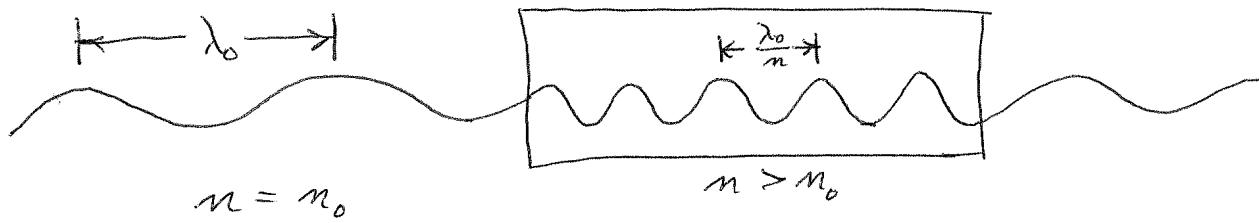
$$I = \frac{c}{8\pi} E_0^2$$

We will generally let $I \sim \underline{E}^* \underline{E}$

k describes the spatial frequency of the light and its direction. It has units of cm^{-1} and essentially gives how many "periods" per cm.

for visible light, $|k| \approx 10^5 \text{ cm}^{-1}$

Refractive Index



typical values

Water $n = 1.33$

glass $n = 1.5$

$$\lambda = \frac{\lambda_0}{n}$$

wavelength
shortened

$$k = \frac{2\pi}{(\frac{\lambda_0}{n})} = \frac{2\pi n}{\lambda_0}$$

spatial frequency
increased

$$c = \frac{c_0}{n}$$

speed
reduced

Complex Refractive Index: Absorption

$$\text{let } n = n + i\epsilon'$$

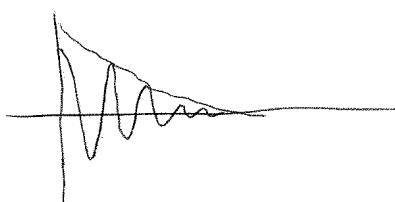
$$k = \frac{2\pi}{\lambda_0} (n + i\epsilon') = \frac{2\pi n}{\lambda_0} + i\epsilon' \quad (\epsilon' = \frac{2\pi}{\lambda_0} \epsilon')$$

then consider a wave

$$E = E_0 e^{i(\omega t + kx)} \Rightarrow E_0 e^{i\omega t} e^{i\frac{2\pi n}{\lambda_0} x} e^{-\epsilon' x}$$

$$= E_0 e^{i(\omega t + \frac{2\pi n}{\lambda_0} x)} e^{-\epsilon' x}$$

\uparrow \uparrow $\underbrace{\qquad\qquad\qquad}_{\text{absorption!}}$
 time oscillator spatial oscillator

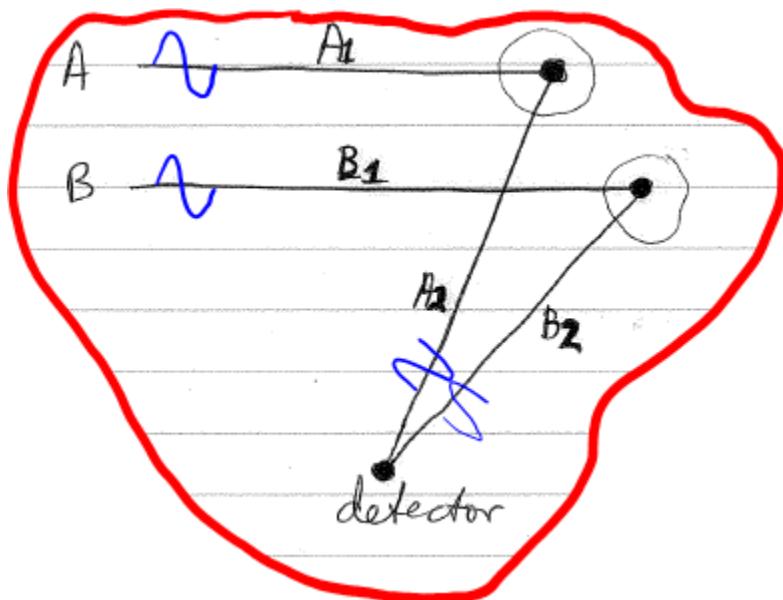


Why stuff scatters.

It's actually more instructive to explain why stuff *doesn't* scatter.¹ Consider the construct below, showing two scattering volumes, A and B, in a sample. These regions aren't chosen at random; rather, they are selected such that the total distance from light source to detector through these two zones differs by a half wavelength (or multiple):

$$A_1 + A_2 = B_1 + B_2 + \lambda/2,$$

where λ is the wavelength. As a result, the waves arrive out of phase, leading to destructive interference...*but only if the scattering strength coming out of the A subvolume matches that coming out of the B subvolume*. If the contents of the two subvolumes are not identical, the destructive interference will be incomplete and scattering will occur.

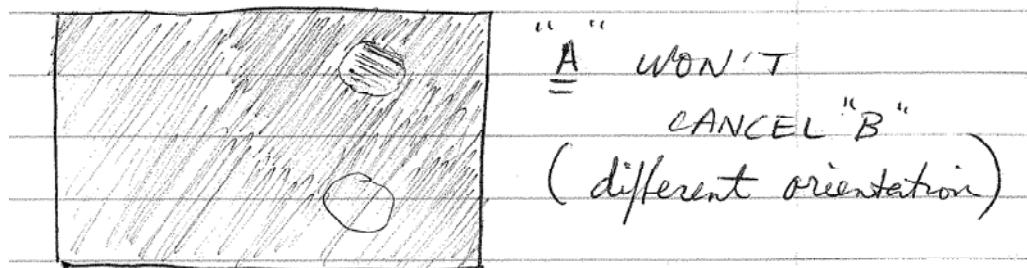
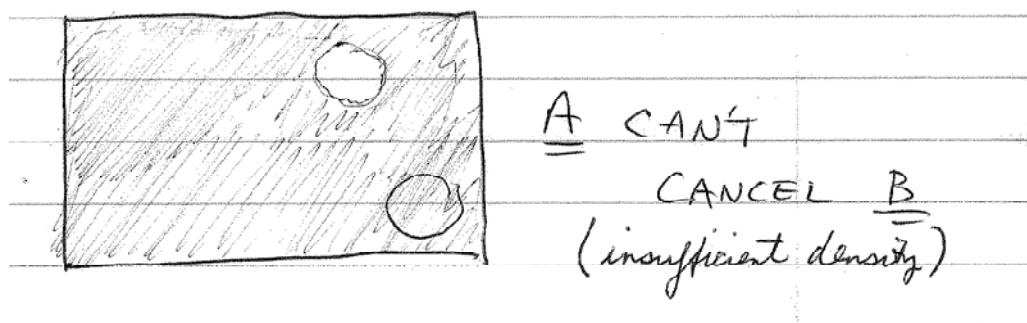
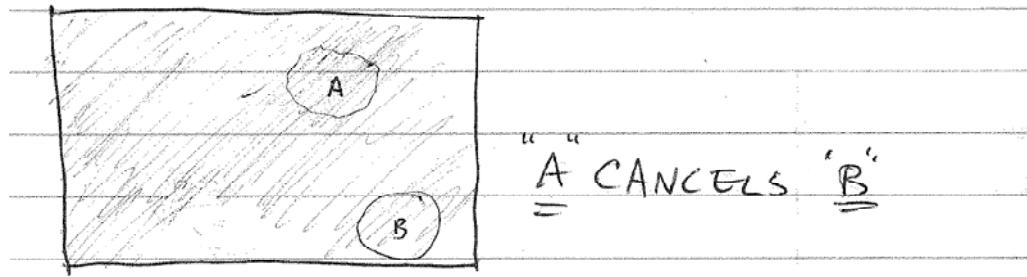


If the subvolumes A and B are identical, scattering cannot occur.

In a sufficiently large sample, any region A has a corresponding region B positioned so that the total scattering will null out if the contents of all pairs are identical. Now we have to consider how the contents can be the same.

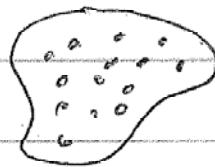
¹ Harkens back to "Slaughterhouse Five" by Ray Bradbury, in which a German Nazi randomly slaps an American prisoner of war. The American asks, "Why me?". His tormentor replies, "Why not you?".

Suppose the contents of A and B contain exactly the same number of identical molecules, randomly positioned and oriented; this would lead to cancellation and no scattering. If the two volumes contain differing numbers, the amplitude coming out of one is insufficient to cancel the scattering derived from the other; while the regions are positioned correctly for destructive interference, the effect is incomplete. Finally, suppose the number of molecules in each subvolume is the same, but their orientation differs. If the orientation is not only geometric but also optical, that could lead to *depolarized* scattering, in which the polarization sense of the outgoing beam differs from that of the incident.



Inhomogeneities lead to scattering.

Let us defer the orientation effects (or is that defects?) for now and consider scattering from systems containing just one component, such as pure nitrogen, pure water or pure carbon (as diamond). Even if each region A and B has identical size, differences in the number of scatterers within can exist due to normal fluctuations. For a sample of population N , the fluctuation δN is estimated as \sqrt{N} . It's easy to see that a 1% fluctuation should be expected if $N = 10,000$.



N particles

$$\delta N \sim \sqrt{N}$$

Relative noise $\frac{\sqrt{N}}{N} = \frac{1}{\sqrt{N}}$

1% Relative noise $\Leftrightarrow N \approx 10,000$

We haven't said anything about the size of the subvolumes, but we should. It's important that the incident light remain in-phase as it crosses the subvolume. A good size, then, might be $\lambda/2$, but let's choose λ for convenience. All our arguments in this section are qualitative anyway. Let's get the pressure of a gaseous system that has a number density, $v = N/V$, corresponding to $10,000/\lambda^3$.

What pressure corresponds to $V = \frac{10,000 \text{ gas atoms}}{\lambda^3}$

$$\lambda = 5000 \text{ \AA} (\text{green-blue})$$

$$PV = NkT$$

$$P = \frac{N}{V} kT = \frac{V kT}{\lambda^3} \quad k = \frac{1.38 \times 10^{-16} \text{ erg}}{\text{mol} \cdot \text{K}}$$

$$T = 300 \text{ K}$$

$$\Rightarrow P = \frac{10,000}{(5000 \times 10^{-8} \text{ cm})^3} \left(1.38 \times 10^{-16} \frac{\text{erg}}{\text{mol} \cdot \text{K}} \right) (300 \text{ K})$$

= Yuck!

Try This instead

$$P = \frac{(10,000 / 6.02 \times 10^{23} \text{ mol})}{(5000 \times 10^{-8} \text{ cm})^3} \left(\frac{82 \text{ mL-atm}}{\text{mol} \cdot \text{K}} \right) (300 \text{ K})$$

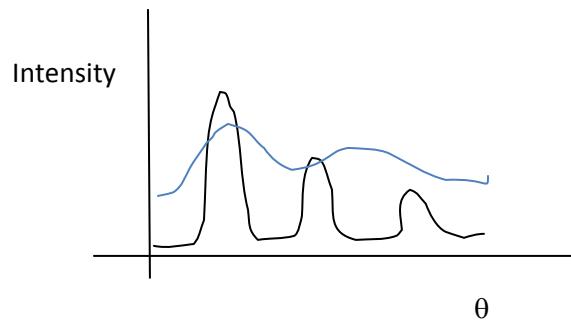
$$\Rightarrow 0.0033 \text{ atm}$$

$$\equiv 2.5 \text{ mm Hg}$$

This pressure, which corresponds to Earth atmosphere at an altitude exceeding 100,000 ft (according to a quick Google search), is already sufficient to cause only 1% fluctuations in the particles within a box of volume λ^3 . On the surface of earth, the fluctuations should be smaller still, yet it turns out even this density not enough to extinguish all the scattering from the particles. One needs even higher number densities. For example, the scattering by water is indeed very low; it is easy to show water has a number density of $v \approx 4 \text{ billion}/\lambda^3$ so $\delta N/N \approx 0.002\%$. Such a low level of fluctuations is considered low enough, and it is often said that:

The scattering of light by a perfect fluid is zero (but no perfect fluid exists).

A perfect crystal, such as diamond also qualifies, but this leads to an interesting point: one cannot get to a density of many scatterers per λ^3 if λ is very small. Water and diamond both scatter x-rays well, but one sees the effect of crystallinity (diamond) or molecular correlation (water) in the scattering pattern, which fluctuates wildly with scattering angle, θ .



Returning to the scattering of visible wavelengths, one usually explains it not as fluctuations $\delta N/N$ but as fluctuations in density. This makes it clear that water will scatter more strongly near its boiling point, for example.

? What about a polymer solution, say $M = 100,000$ at $c = 1 \text{ mg/mL}$. How many scatterers in a volume of λ^3 ? Then?

① First, let us agree that the number of solvent molecules is really huge. There will be some scattering due to density fluctuations, but it should be small (greater than air, due to the huge number of particles, but still very small on a per particle basis).

② Convert the concentration

$$\left(10^3 \frac{\text{g}}{\text{mL}}\right) \left(5000 \times 10^{-8} \text{ cm}\right)^3 \left(\frac{6 \times 10^{23} \text{ particles}}{100,000 \text{ g}}\right) = 750$$

Two things: 1) The per particle scattering of a polymer

can be very large

$$2) \frac{\sqrt{750}}{750} \times 100 \approx 3.5\% - \text{significant fluctuations}$$

Mead

We should expect an easy time measuring scattering from polymer solutions, but despite their high mass, one has to reach very high dilutions to measure single particles at once. If you see in the literature a study that claims to have been performed in the "single particle limit", it means something else—e.g., the particles are not interacting, though many are normally detected in the instrument.

Summary so far:

- It's not surprising that objects with a high density don't scatter much at visible wavelengths.
- We should expect polymers to scatter a bit better.

Before moving on, let's look at our four questions and/or observations:

1. We haven't explained why big particles scatter more than small ones yet.
2. We haven't explained angular dependence yet.
3. We haven't explained why the sky is blue or sunsets are red, nor have we explained polarization or the reason why the sky seems distant. Nevertheless, from the above consideration it should seem plausible that many miles of atmosphere could scatter light. Also, at the top of the atmosphere, the cancellation effect will be weak (but there are few scatterers to begin with, too).
4. We haven't explained diffraction...but we have pointed out that wild oscillations will occur when the interparticle distance approaches the wavelength. This is true of solids *and liquids* but the oscillations are less profound in the latter case.

It might seem like little progress has been made, but you are beginning to think like a scatterbrain.

Scattering from a single, small particle in a gas.

References: Tanford p.278

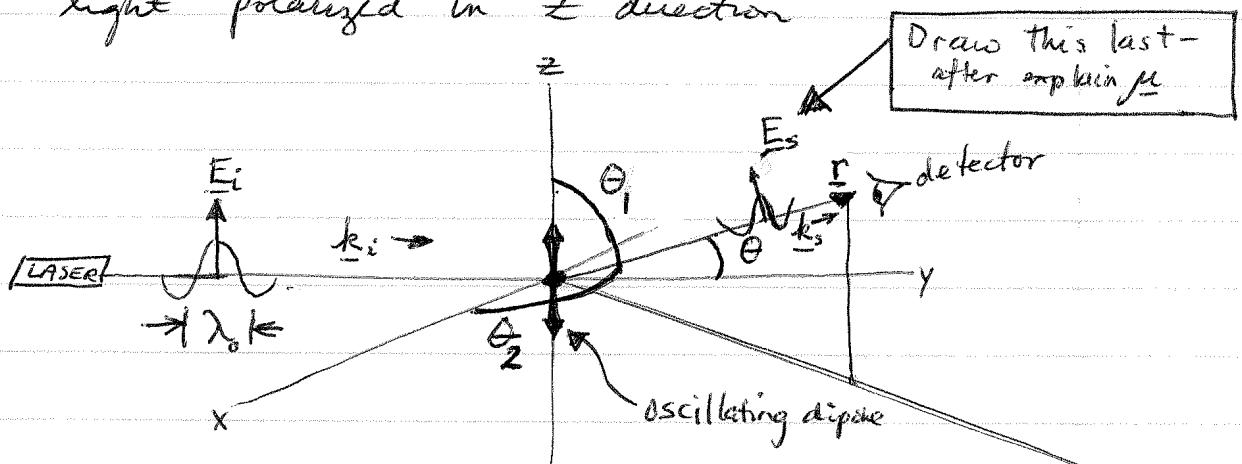
Van Holde Ch. 9

Zimmon, Stein & Doty Polym. Bull. 1 (415)

90-119 (1948)

Johnson & Gabriel, Laser LS, Dover, Mineola NY
(1981)

- Consider a gas atom (size $\ll \lambda$) located at origin
- light incident along y axis
- light polarized in z direction



let \underline{k}_s = vector || scattered light, magnitude $\frac{2\pi}{\lambda_0}$.

let \underline{k}_i = vector || incident light, magnitude $\frac{2\pi}{\lambda_0}$.

let θ_1 = The polar angle (what we often call ϕ or θ)

θ_2 = angle between x axis & scattering direction

θ = "the" scattering angle = angle btwn. \underline{k}_i & \underline{k}_s

r = vector btwn. scatterer & detector

μ = induced dipole at origin, as electrons respond to the incident field E_i

DK - That's the set-up. Now what happens?

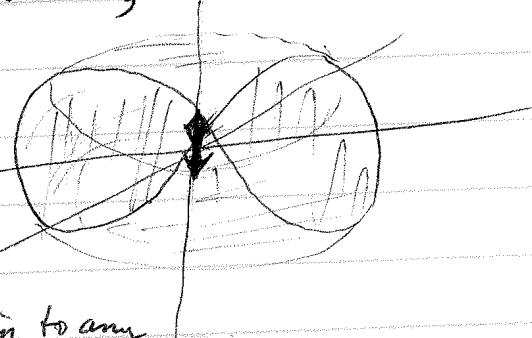
- \underline{E}_i is a rapidly oscillating E & H field
 - The E field is \parallel to z direction. The magnitude of \underline{E}_i varies at $\sim 10^{14}$ - 10^{15} Hz.
 - Inside the atom, the electrons & protons dance to the tune of the incident light.
 - To first approximation, the electrons do most of the moving (proton is 1836 times heavier).
- [Born Approximation]

$$\mu = \alpha \underline{E}_i \quad \left\{ \begin{array}{l} \text{assume } \mu \parallel \underline{E}_i \\ \text{This actually isn't true} \end{array} \right.$$

Units: $(\text{esu}\cdot\text{cm}) = \alpha \text{ units } (\text{esu}\cdot\text{cm}^{-2})$
 $\alpha \text{ units are } \text{cm}^3 \quad (\alpha \text{ related to volume})$

- An oscillating dipole re-radiates light. This is how antennae work. Now the intensity radiated by an oscillating dipole has a well known pattern, like a donut with no hole in it.

Take a figure 8 on its side & rotate it around z axis.



The distance from origin to any point on donut represents the intensity in that direction.

$$\cos(90 - \theta_i) = \sin(\theta)$$

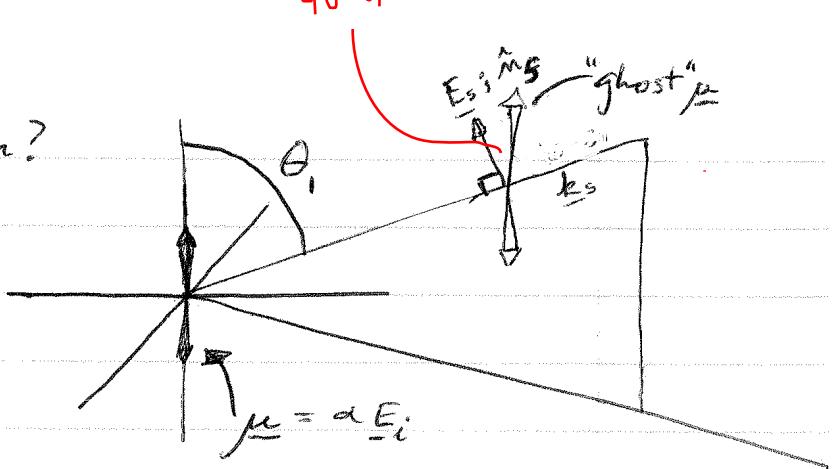
LS-5

Why the donut pattern?

translate the μ dipole

to some remote point

(call this "ghost" μ)



- It is a property of light that the propagation direction is exactly \perp to the electric field. So E_s is \perp to k_s .
- The amplitude of E_s depends on θ_i .
- From the drawing, it can be seen that $|E_s| \sim \sin \theta_i$.
- So to get the most scattering, operate in the x, y plane.

Now we can do the math.

At origin $E_i = \hat{z} E_0 e^{i(\omega t)}$ $\omega = 2\pi c/\lambda$

: will be directed as shown. Call this vector \vec{E}_i

E_s : will be proportional to $\sin \theta_i$,

: will be proportional to $\left(\frac{d^2 \mu}{dt^2}\right) \left(\frac{1}{rc^2}\right)$

{ Why? because $\frac{d\mu}{dt}$ is a current, and
accelerating current gives rise to radiation
The rc^2 term ensures intensity drops as r^{-2}
The c^2 part is dimensionally correct at least }

: will be modified by a phase factor $k_s \cdot \frac{1}{r}$

by the time we measure it at the detector.

$$E_s = \frac{\hat{z}}{rc^2} \frac{d^2 \mu}{dt^2} = \left(\frac{d^2}{dt^2} \left[\alpha E_0 e^{i(\omega t - k_s \cdot \frac{1}{r})} \sin \theta_i \right] \right) \frac{1}{M_s} \left(\frac{1}{rc^2} \right)$$

- The two derivatives bring down $(i\omega)^2 = -\omega^2$
- Everything else comes back as it was, since $\frac{d}{dx} e^x = e^x$

Thus $E_s = \frac{\alpha}{rc^2} E_0 (-\omega^2) e^{i\omega t - iks \cdot \hat{r}} \sin \theta_i$

Useful to dissect this

The direction \perp to \hat{k}_S in which E field lies

incident light strength

Polarizability = Scattering Strength

field fades as $\frac{1}{r}$ from $E \& M$ (some squared velocity is needed by dimensions)

$\sin \theta_i$ oscillations

spatial phase term

geometry

Scattered light field goes up as square of frequency - comes from $E \& M$: radiation is acceleration of current.

- sign says that there is a 180° phase change on scattering.

Instantaneous Intensity is actually $\sim |E \times H|$ but we do OK to compute $E^* E$ which gives I_{average}.

$$I_s = E_s^* E_s = \frac{1}{r^2 c^4} \alpha^2 E_0^2 \omega^4 \sin^2 \theta_i$$

$$\left. \begin{aligned} v &= c/\lambda_0 \\ \omega &= 2\pi\nu = 2\pi c/\lambda_0 \\ (\omega/c)^4 &= \left(\frac{2\pi}{\lambda_0}\right)^4 = \frac{16\pi^4}{\lambda_0^4} \end{aligned} \right\}$$

$$I_s = \frac{\alpha^2 (16\pi^4)}{r^2 (\lambda_0^4)} \sin^2 \theta_i$$

geometry

$I_s \sim (\text{volume})^2$

blue scatters more than red

fades like r^2

$$\left. \begin{aligned} \text{let } E_0^2 &= I_i \\ E_s^* E_s &= I_i \end{aligned} \right\}$$

Note: the λ_0^4 in this expression came via frequency terms. It would still be λ_0^4 even in a general medium of $RI^* = n$. However, for large particles, we will see that λ_0/n also enters—not here in the fundamental scattering power, but rather in the angular dependence.

*($RI =$ refractive index).

Still Single Particle Dilute Gas, but now convert from α to n .

The Debye relation:

$$n^2 - 1 = 4\pi V \alpha$$

$\left. \begin{array}{l} \text{esu units} \\ \text{gas phase} \\ V = \# \text{ density} \end{array} \right\}$

But for dilute gas, $n^2 \approx 1$, so expand:

$$n^2(c) = \underbrace{n^2(c=0)}_1 + 2n \left(\frac{\partial n}{\partial c} \right)_{c=0} c + \dots \quad \left. \begin{array}{l} c = \text{conc. of} \\ \text{gas ass} \\ g/mL \end{array} \right\}$$

$$\Rightarrow n^2(c) - 1 = 2 \cdot \cancel{1} \cdot \left(\frac{\partial n}{\partial c} \right)_{c=0} c$$

or... $n^2 - 1 = 2c \left(\frac{\partial n}{\partial c} \right)$

$\left. \begin{array}{l} \text{drop explicit } c=0 \\ \text{reminders} \end{array} \right\}$

Comparing the 2 boxed equations: $\alpha = \frac{2c(\partial n / \partial c)}{4\pi V}$

But $V = \frac{M}{N}$, so: $\alpha = \frac{2M(\partial n / \partial c)}{4\pi N}$

$$\boxed{\alpha = \frac{M(\partial n / \partial c)}{2\pi N}}$$

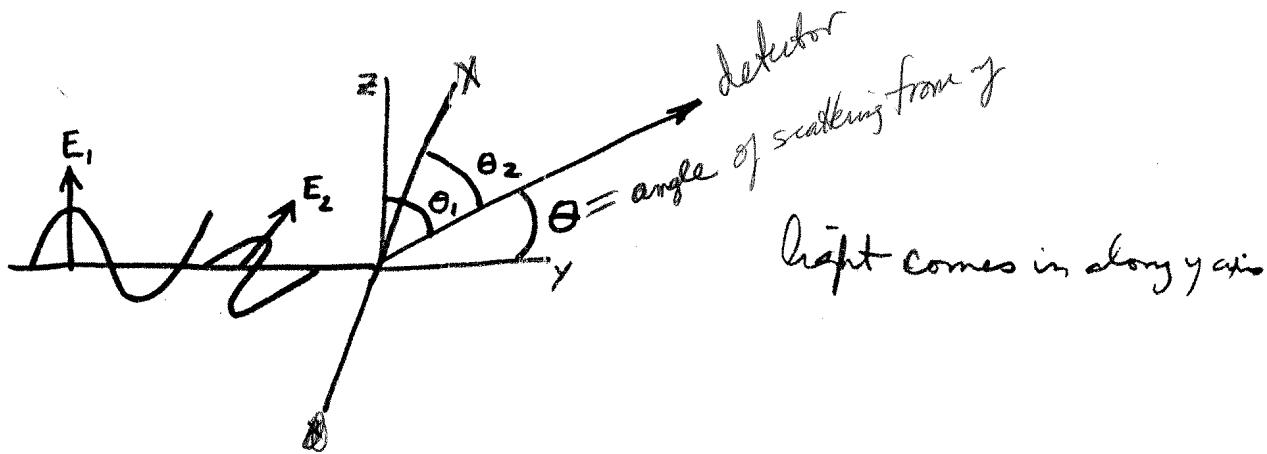
Our equation was $\frac{I_s}{I_i} = \frac{\alpha^2}{r^2} \left(\frac{16\pi^4}{\lambda_0^4} \right) \sin^2 \theta,$

With expression for α we get:

$$\boxed{\frac{I_s}{I_i} = \frac{4\pi^2}{\lambda_0^4 r^2} \left(\frac{M}{N} \right)^2 \left(\frac{\partial n}{\partial c} \right)^2 \sin^2 \theta,}$$

Scattering from one gas particle

Extension to Unpolarized Light



Unpolarized light = $E_1 \parallel$ to Z direction plus
 $E_2 \parallel X$ direction where there is arbitrary phase
 relationship ~~to~~ between E_1 and E_2 . Both E_1 and
 E_2 have same intensity.

Light Scattering by Particles

$$\frac{I_{1,S}}{\frac{1}{2} I_i} = \frac{4\pi^2 \sin^2 \theta_1 \left(\frac{dn}{dc}\right)^2 M^2}{N^2 r^2}$$

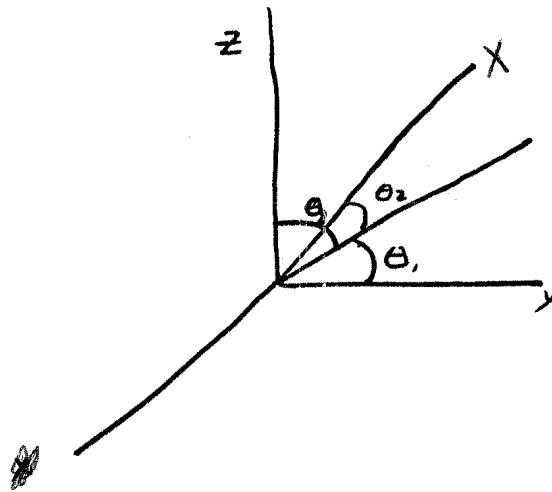
$$\frac{I_{2,S}}{\frac{1}{2} I_i} = \frac{4\pi^2 \sin^2 \theta_2 \left(\frac{dn}{dc}\right)^2 M^2}{N^2 r^2}$$

$$\Rightarrow I_S(\theta) = \frac{2\pi^2 \left(\frac{dn}{dc}\right)^2 M^2}{N^2 r^2} (\sin^2 \theta_1 + \sin^2 \theta_2)$$

↑ →
 Unpolarized light will be indicated by $I_S(\theta)$ rather than I_S

$$\text{But } \sin^2 \theta_1 + \sin^2 \theta_2 = 1 + \cos^2 \theta$$

Why :



Because $\theta_1, \theta_2, \theta$ are direction angles we know
that :

$$\cos^2 \theta_1 + \cos^2 \theta_2 + \cos^2 \theta = 1$$

$$\left(\text{from } x^2 + y^2 + z^2 = r^2 \right. \\ \left. \Rightarrow \left(\frac{x}{r}\right)^2 + \left(\frac{y}{r}\right)^2 + \left(\frac{z}{r}\right)^2 = 1 \right)$$

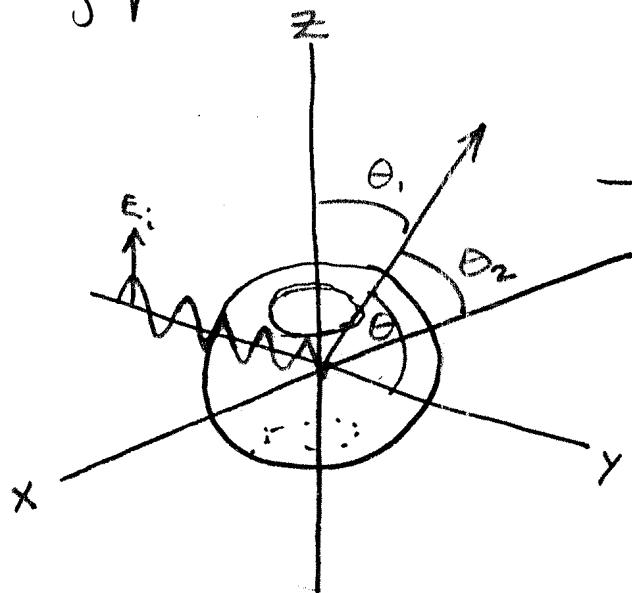
$$\Rightarrow 1 - \sin^2 \theta_1 + 1 - \sin^2 \theta_2 + \cos^2 \theta = 1$$

$$\Rightarrow \sin^2 \theta_1 + \sin^2 \theta_2 = 1 + \cos^2 \theta \quad \text{QED}$$

$$\therefore \frac{I_s(\theta)}{I_i} = \frac{2\pi^2 \left(\frac{dn}{dc} \right)^2 M^2}{N^2 R^2} (1 + \cos^2 \theta)$$

This produces a different dependence on angle than for planarized.

Summary of Rayleigh

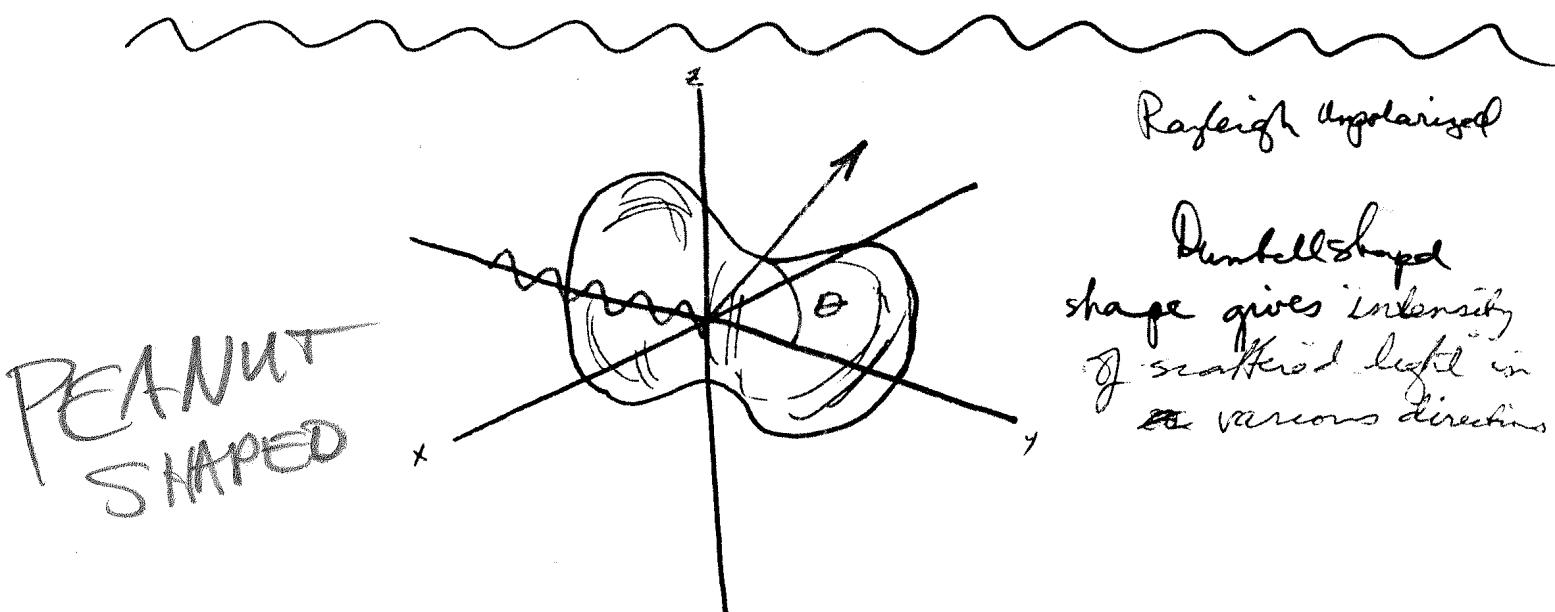


Polarized Rayleigh

Donut shaped figure represents the intensity of the scattered light in various directions.

At const θ_1 , $\theta_2 \leq \theta$ can be varied anywhere

$$\frac{I_s}{I_i} = \frac{4\pi^2 \sin^2 \theta_1 M^2 (\lambda^2/c)^2}{\lambda^4 r^2 N^2}$$



Rayleigh unpolarized

Dumbbell shaped shape gives intensity of scattered light in various directions.

$$\frac{I_s(\theta)}{I_i} = \frac{2\pi^2 M^2 (\lambda^2/c)^2}{\lambda^4 r^2 N^2} (1 + \cos^2 \theta)$$

More Than One Particle: Delicate Gas

Now, electric field at detector will be sum of many terms:

$$\underline{E}_{\text{detector}} = \underline{E}_1 + \underline{E}_2 + \underline{E}_3 + \dots + \underline{E}_n$$

$$I_{\text{detector}} \propto \left(\underline{E}_1^* \underline{E}_1 + \underline{E}_1^* \underline{E}_2 + \dots + \underline{E}_1^* \underline{E}_n \right) \cdot \left(\underline{E}_n^* \underline{E}_1 + \dots + \underline{E}_n^* \underline{E}_n \right)$$

Remember: Instantaneously,
This isn't really true!

$I \propto \underline{E} \times \underline{H}$, but letting it be
 $\underline{E}^* \underline{E}$ will suit our purposes
for time averaged intensity

$$E_i = E_0 \sin(\omega t + \phi_i) = E_0 e^{i(\omega t + \phi_i)}$$

↗ phase factor giving
position of particle
w.r.t. source and detector.

For unstructured gas, off-diagonal terms go off to zero, because phase is random. \Rightarrow Only diagonal terms survive.

Averaging over time gives:

$$I_{\text{detector}} \propto n E_0^2 \quad \left\{ \begin{array}{l} \text{Note similarity to} \\ \langle r^2 \rangle = nl^2 \end{array} \right\}$$

Thus $I(N \text{ molecules in a gas}) = N I(\text{one molecule in a gas})$

$$I_s = 2\pi^2 NM^2 (\frac{\partial n}{\partial c})^2$$

$$I_i = \lambda_0^4 r^2 N^2$$

G

↑ geometry factor

$$G = 2 \sin^2 \theta, \text{ Vert. Pol.}$$

$$\text{or } (1 + \cos^2 \theta) \text{ Random Pol.}$$

1) ÷ both sides by V = scattering volume

$$2) \frac{NM}{VV} = \frac{g_2}{V} = c$$

$$\frac{I_s}{I_i} = \frac{2\pi^2 c M (\frac{\partial n}{\partial c})^2}{\lambda_0^4 r^2 N} G$$

dilute, multiparticle gas

$$G = 2 \text{ for usual arrangement } \theta_i = 90^\circ$$

$$I_s \sim c M = m M^2 = m R^6$$

Two-component solution: small particles
at low concentration.

For gas, we had:

$$n^2 - 1 = 2c \left(\frac{dn}{dc} \right)$$

If we examine the excess scattering, this becomes:

$$n^2 - n_0^2 = 2n_0 c \left(\frac{dn}{dc} \right)$$

$$\therefore \frac{I_s}{\sqrt{I_i}} = \frac{2\pi^2 n_0^2 \left(\frac{dn}{dc} \right)^2 M c}{N \lambda_0^4 R^2} G(\theta)$$



geometry factor
 $2 \sin^2 \theta$, (polarized)
 $(1 + \cos^2 \theta)$ (unpolarized)

Molecular Weight

We will soon see that the equation must be modified for:

- 1) Thermodynamic interaction
- 2) Finite size.

But it's a good time to see what the equation can do in the ideal gas limit.

The equation says $I_s \sim c M$

what average do we get for a polydisperse gas?

$$I_s \sim \sum c_i M_i = c_{\text{total}} M_{LS}$$

↑
define

$$\text{Thus, } M_{LS} = \frac{\sum c_i M_i}{c_{\text{total}}} = \frac{\sum \left(\frac{g_i}{v}\right) M_i}{\left(\frac{\sum g_i}{v}\right)}$$

$$M_{LS} = \sum w_i M_i = M_w$$

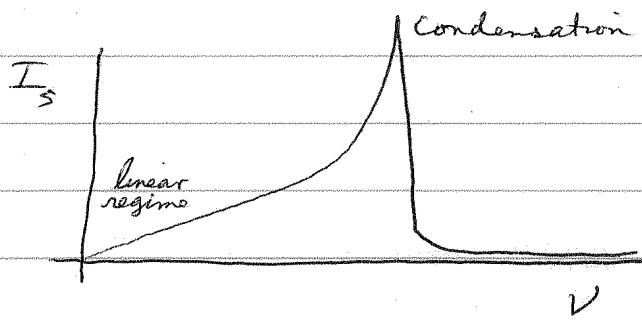
LS gives weight average mass!

This result is for dilute solution limit, but later we shall see how we can always extrapolate to this limit.

Thermodynamic Approach

It isn't quite as deep as wave-particle duality in quantum mechanics, but LS has a dual nature, too: sometimes, a particle approach is best, while at other times it may be better to adopt a thermodynamic viewpoint. The thermodynamic picture is more general and should reduce to the particle picture in the dilute limit.

To begin with, consider I_S vs. V where V is the number density of small particles, say H_2O particles in air.

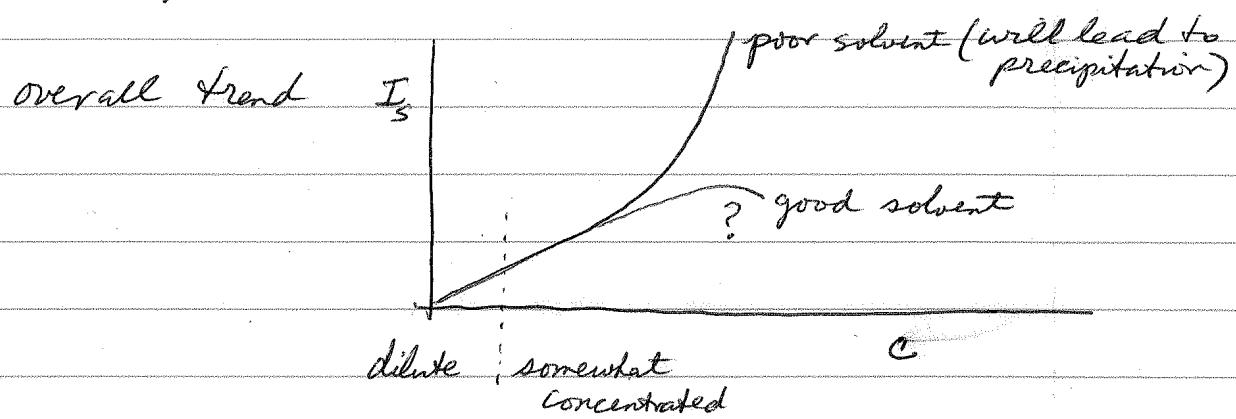


The maximum scattering is when the particles have a size $\approx \lambda$. Particles larger than that do scatter a lot, but only to the forward direction. This is why one can see through rain but not fog. (R. Stein, F2000).

At low number density, there is a linear increase, as we have discussed by "matrix-like" summation on p. LS-12.

Let us now recast this as a polymer/solute problem rather than water/air.

Two-component, non-dilute solutions (Tanford)



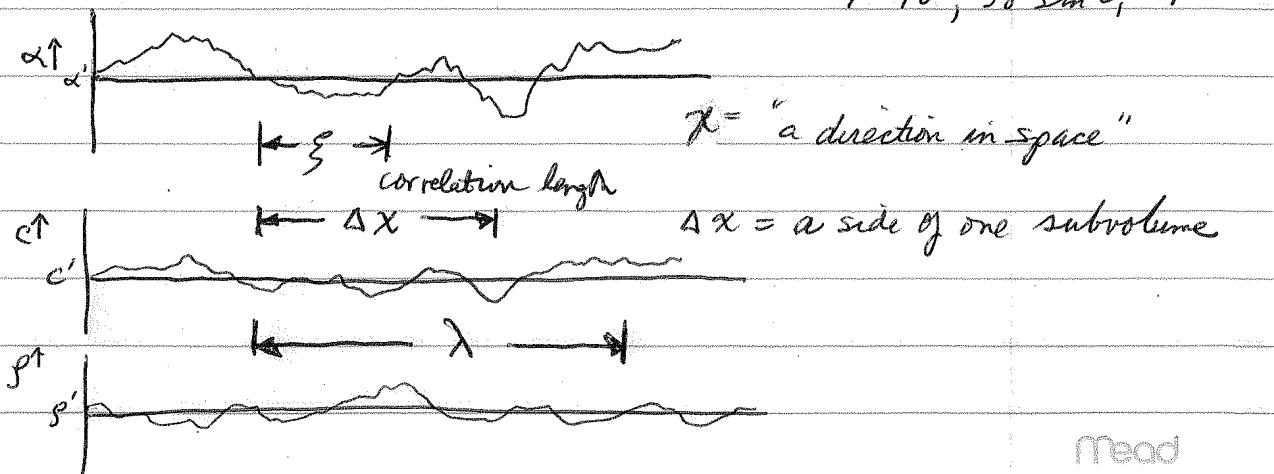
The turnaround in good solvents is implicit in that a pure, crud-free, glassy polymer (like PMMA) scatters little.

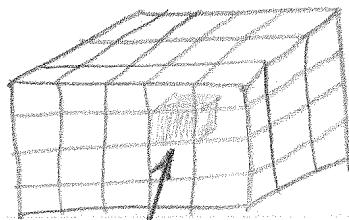
dilute: we can adapt from gas limit
somewhat concentrated: we need thermo, and it should behave well in the dilute limit, too.

To do Thermo approach, we go all the way back to:

$$\frac{I_s}{I_0} = \frac{16\pi^4 \alpha^2 \sin^2 \theta}{\lambda^4 R^2} \rightarrow \text{let's stipulate that } \theta_i = 90^\circ, \text{ so } \sin^2 \theta_i = 1$$

3 PLOTS





m volume elements

$$\Delta x \quad \text{Volume} = \gamma = (\Delta x)^3$$

→ We want scattering from a single volume element - like single particle case but scat. arises out of fluctuations.

→ We must consider how to add these together.

* $\gamma = (\Delta x)^3$ must be large enough to constitute a thermodynamic system.

* We want the interaction length ξ to be $\leq \Delta x$. Then $\delta \alpha_i$ in one volume element and $\delta \alpha_j$ in another are unconnected. The product $\delta \alpha_i \delta \alpha_j$ could be + or -.

* But Δx must be less than λ , so that each volume can be treated as a point source.

i.e. $[\xi < \Delta x < \lambda]$ or actually $\frac{2\pi}{\xi}$ probably better gauge than λ .

These are tough demands. I'm afraid I just don't know.

- Cannot treat phase separating systems since $\xi \rightarrow \infty$

May be problems with v. large molecules, larger λ , possibly worst in really dilute solutions?

You might write

$$\frac{I_s}{I_0} = \frac{16\pi^4}{\lambda_0^4 r^2} (\alpha' + \delta\alpha_i)^2$$

\equiv intrinsic scattering power of
one isolated volume element, i.e.

What will the scattering be from all of the elements?

$$E_s = (\alpha' + \delta\alpha_1) e^{\phi_1} + (\alpha' + \delta\alpha_2) e^{\phi_2} + \dots$$

$$\Rightarrow I_s = \left((\alpha' + \delta\alpha_1)^2 + (\alpha' + \delta\alpha_1)(\alpha' + \delta\alpha_2) e^{\phi_2 - \phi_1} + \dots + \right. \\ \left. (\alpha' + \delta\alpha_1)(\alpha' + \delta\alpha_2) e^{\phi_2 - \phi_1} + (\alpha' + \delta\alpha_2)^2 + \dots + \right)$$

Although the formalism is by now familiar, this is different. The phase factors are not random, as when the indices on them represented mobile gas atoms. They are fixed, representing the location of a pair of subvolumes.

In general, this won't be easy to simplify. However, if the correlation length is not too long, we stand a good chance of the fluctuation $\delta\alpha_i$ being uncorrelated from $\delta\alpha_j$. In this case, maybe the excess scattering can be estimated from the diagonals

Somewhat simpler case 1

If long-range interactions are weak, off diagonals go to zero, at least insofar as the $\delta\alpha$ effects are concerned.

$$I_S \sim \sum_{i=1}^m (\alpha_i + \delta\alpha_i)^2 = \sum_{i=1}^m (\alpha^2 + 2\delta\alpha_i + (\delta\alpha_i)^2)$$

$\sum \alpha^2$ term We could take this one & subdivide very finely i.e. right down to the angstrom level because α is the same in any volume element, no matter what its size. This term is liquid-like and goes to zero. Keeping it would only represent an artifact of the coarse-grained subdivision of volume elements that we have used in our matrix and in computing the phase factors ϕ .

$\sum_i 2\delta\alpha_i$ term $\Rightarrow 0$ because $\sum_i \delta\alpha_i = 0$

The only term that contributes is $\sum (8\alpha)^2 = m \overline{\delta\alpha^2}$

$$I_S \sim m \overline{\delta\alpha^2}$$

Now... how to get $\overline{\delta\alpha^2}$?

$$\text{let } \overline{\delta\alpha} = \underbrace{\delta\alpha_s}_{\substack{\text{density fluctuations} \\ \text{occur for both solution} \\ \text{and pure solvent}}} + \underbrace{\delta\alpha_c}_{\substack{\text{concentration fluctuations} \\ \text{don't occur for} \\ \text{pure solvent.}}}$$

$$\overline{\delta\alpha^2} = \overline{\delta\alpha_s^2} + 2\overline{\delta\alpha_s}\overline{\delta\alpha_c} + \overline{\delta\alpha_c^2}$$

↓ ↓

Assume $\overline{\delta\alpha_s \delta\alpha_c}$ is uncoupled $\Rightarrow \text{average} = 0$

Assume to be

The same for solvent
and solution. So if we talk about excess intensity
scattered above solvent then we are OK to ignore this term.

$$\text{Then } \overline{\delta\alpha^2} = \overline{\delta\alpha_c^2}$$

↓

2 assumptions

We will now develop $\overline{\delta\alpha_c^2}$

Let us regroup some: We have shown

$$I_s \sim m \overline{\delta\alpha_c^2}$$

\nearrow
excess
above
solvent $\rho < 1x < \lambda$ $m = \# \text{ volume elements}$

Plug into the multiple particles random phase expression:

$$\frac{1}{V} \frac{I_s}{I_i} = \frac{m}{V} \left(\frac{16\pi^4}{\lambda_0^4 r^2} \sin^2 \theta_j \right) \overline{\delta\alpha_c^2}$$

\uparrow assumes V polarization

note: $m = \frac{1}{V} = \frac{1}{(\lambda k)^3}$

$$\frac{1}{V} \frac{I_s}{I_i} = \frac{1}{V} \left(\frac{16\pi^4}{\lambda_0^4 r^2} \sin^2 \theta_j \right) \overline{\delta\alpha_c^2}$$

Rescale, add \int

As for gases, we must get from α to c .

$$\text{Recall for gases: } n^2 - 1 = 2c(\frac{dn}{dc}) \quad \left[\begin{array}{l} \text{expansion of} \\ n^2 = n_0^2 + 2n_0(\frac{dn}{dc}) \end{array} \right]$$

$$n^2 - 1 = 4\pi V \alpha \quad \text{Clausius-Mosotti Eqn.}$$

$$\Rightarrow \alpha = \frac{c(\frac{dn}{dc})}{2\pi V}$$

$$\text{For polymer solutions: } n^2 - n_0^2 = 2n_0 \frac{\delta c}{c} (\frac{dn}{dc}) \quad \left[\begin{array}{l} \text{expand about} \\ \text{pure solvent} \end{array} \right]$$

$$\text{eqt} \quad n^2 - \frac{n_0^2}{c} = \frac{4\pi V \frac{d\alpha}{dc}}{f}$$

Assuming $n_c \approx n_0$ then equate A & B assuming f terms vanish for the reasons discussed

$$\frac{\delta c}{c} = \frac{n_0(\frac{dn}{dc})}{2\pi V}$$

$$\text{or } \frac{(d\alpha)^2}{4\pi^2 V^2} = \frac{n_0^2(\frac{dn}{dc})^2}{(dc)^2}$$

$$\text{But } V = m = \frac{1}{\rho}$$

$$\therefore \frac{(d\alpha)^2}{4\pi^2} = \frac{n_0^2 \rho^2 (\frac{dn}{dc})^2}{(dc)^2}$$

Now put this into eqn. from previous page.
One power of π cancels. A $4\pi^2$ cancels, too.

$$\frac{S_{\text{over}}}{\sqrt{I_0}} = \frac{4\pi^2 \Psi M_0^2 (\delta n/\delta c)^2}{\lambda_0^4 r^2} \overline{(\delta c)^2}$$

Same as

Tanford q. 17-18

except we limited
to $\theta_i = 90^\circ$

Now we require an expression for $\overline{(\delta c)^2}$, the mean square concentration fluctuation. This is going to come from Boltzmann statistics.

$$P(\delta c) = e^{-\Delta G(\delta c)/kT}$$

density

= probability of an element being at a concentration $c' + \delta c$.

where ΔG is the free energy difference between concentrations $c' + \delta c$ and c' on a per molecule basis.

This $\Delta G(\delta c)$ will come from an expansion:

$$\Delta G(\delta c) = G(c' + \delta c) - G(c') = \left(\frac{\partial G}{\partial c}\right)_c \delta c + \frac{1}{2!} \left(\frac{\partial^2 G}{\partial c^2}\right)(\delta c)^2 + \dots$$

But c' is the equilibrium concentration. By

definition $\left(\frac{\partial G}{\partial c}\right)_{c'} = 0$.

$$\therefore \Delta G(\delta c) \approx \frac{1}{2} \left(\frac{\partial^2 G}{\partial c^2}\right)_{c'} (\delta c)^2$$

$$\text{and } P(\delta c) = e^{-\frac{1}{2} \left(\frac{\partial^2 G}{\partial c^2}\right)_{c'} (\delta c)^2 / kT}$$

just a constant
 call it A

NOT A BAD PLACE
TO CONTEMPLATE
WHAT HAPPENS
WHEN $\frac{\partial^2 G}{\partial c^2} > 0$

$$\therefore \frac{1}{(\delta c)^2} = \frac{-A(\delta c)^2 d(\delta c)}{\int_0^\infty e^{-A(\delta c)^2} d(\delta c)}$$

actually from $-\infty$ to ∞ but
 does not really require it
 because (δc) is squared

$$\left\{ \text{Integrals of type } \int_0^\infty x^2 e^{-Ax^2} dx \right\}$$

and $\int_0^\infty e^{-Ax^2} dx$

The result of evaluating these standard integrals is:

$$\overline{(\delta c)^2} = \frac{kT}{\left(\frac{\partial^2 G}{\partial c^2}\right)_{T,p}} \Big|_{c=c'}$$

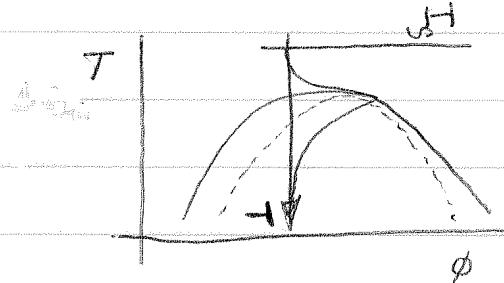
Henceforth, it is understood that derivatives are evaluated at c' .

Does it make sense? Well, G is within a constant of ΔG_{mix} and we saw that at $\frac{\partial^2 \Delta G_{mix}}{\partial c^2} = 0$ we have

the SPINODAL CURVE, which means that B16 concentration fluctuations occur without hesitation.

$\overline{(\delta c)^2}$ does not really go off to ∞ because the terms we truncated prevent that.

So, in fact, The scattering just gets real large.



I vs. T is actually used to locate spinodals, in a method called PICS: pulse induced critical scattering.

Now we must calculate $(\partial^2 G / \partial c^2)$.

Let n_1 and n_2 be mole #'s such that *

$$n_1 \bar{V}_1 + n_2 \bar{V}_2 = \gamma$$

\uparrow \uparrow

Tanford claims $\bar{V}_1 = \bar{V}_2 (c')$.

Not true, I think, but a close approximation.

$$\Rightarrow dn_1 = -dn_2 \frac{\bar{V}_2}{\bar{V}_1} \text{ since } \gamma = \text{constant}$$

$$\text{Now } dG = \mu_1 dn_1 + \mu_2 dn_2 @ \text{const } T, P$$

$$dG = \left(\mu_1 \left(-\frac{\bar{V}_2}{\bar{V}_1} \right) + \mu_2 \right) dn_2 \quad \boxed{\text{EQ. A}}$$

$$\text{But } \frac{n_2}{\gamma} = \frac{c}{M} \quad \text{That is... mol} = \frac{\left(\frac{g}{m}\right)}{mL} \quad \frac{\left(\frac{g}{mol}\right)}{(mL)}$$

$$\text{So... } dn_2 = \frac{\gamma}{M} dc \quad \boxed{\text{EQ. B}}$$

$$\text{Plug B into A} \Rightarrow dG = \left(\mu_1 \left(-\frac{\bar{V}_2}{\bar{V}_1} \right) + \mu_2 \right) \left(\frac{\gamma}{M} \right) dc$$

$$\text{So... } \frac{\partial G}{\partial c} = \left(\mu_2 - \frac{\bar{V}_2}{\bar{V}_1} \mu_1 \right) \left(\frac{\gamma}{M} \right)$$

* It can be argued (Dr. Carpenter) that if γ is fixed and T, P are also, is not the system then fixed? Answer: No! Because composition can still fluctuate.

So... take 2nd deriv to get:

$$\left(\frac{\partial^2 G}{\partial c^2}\right)_{T_p} = \left(\frac{\partial \mu_2}{\partial c} - \frac{V_2}{V_1} \frac{\partial \mu_1}{\partial c}\right) \frac{\psi}{M}$$

↑ we have assumed $\frac{\partial}{\partial c} \left(\frac{V_2}{V_1}\right) = 0$

Now apply Gibbs-Duhem: $d\mu_2 = -\frac{n_1}{n_2} d\mu_1$

$$\left(\frac{\partial^2 G}{\partial c^2}\right)_{T_p} = \underbrace{\left(-\frac{n_1}{n_2} - \frac{V_2}{V_1}\right)}_{\text{↑ appears to be no real physical significance to this term, but it does simplify.}} \left(\frac{\partial \mu_1}{\partial c}\right) \frac{\psi}{M}$$

↑ appears to be no real physical significance to this term, but it does simplify.

$$\left(\frac{\partial^2 G}{\partial c^2}\right)_{T_p} = -\left(\frac{m_1 V_1 + m_2 V_2}{m_2 M M}\right) \left(\frac{\partial \mu_1}{\partial c}\right) \frac{\psi}{V_1}$$

$$= -\left(\frac{1}{c}\right) \left(\frac{\partial \mu_1}{\partial c}\right) \frac{\psi}{V_1}$$

Thus, $(\delta c)^2 = \frac{kT}{\left(\frac{\partial^2 G}{\partial c^2}\right)_{T_p}} \Rightarrow \frac{-kT c V_1}{\psi (\delta \mu / \delta c)}$

But we had $\frac{1}{V} \frac{I_s}{I_i} = \frac{4\pi^2 \Psi_m^2 (\delta \eta / \delta c)^2}{\lambda_o^4 r^2} \frac{(\delta c)^2}{(\delta c)^2}$

So this evolves to:

$$\frac{I_s}{V} = \left(\frac{4\pi^2 \Psi m_e (\partial \mu / \partial c)^2}{\lambda_0^4 r^2} \right) \times \left(\frac{-kTc\bar{V}_i}{\Psi (\partial \mu / \partial c)} \right)$$

Note: Ψ cancels out, as it must because we defined it sort of vaguely BUT to meet specific criteria. So we needed it ... and we needed to get rid of it!

$$\frac{I_s}{V} = \frac{4\pi^2 m_e (\partial \mu / \partial c)^2}{\lambda_0^4 r^2} \frac{c}{\left(\frac{-1}{kT\bar{V}_i} \right) (\partial \mu / \partial c)}$$

But what is $\partial \mu / \partial c$? It clearly represents the energy to generate a concentration fluctuation. If it takes lots of energy to make a fluctuation, then they will not occur... and scattering declines.

Well, we have a formalism for all this already.

$$-\pi V_i^\circ \approx \mu_i - \mu_i^\circ = -RT\bar{V}_i c \left(\frac{1}{M} + A_2 c + A_3 c^2 + \dots \right)$$

$$\Rightarrow \frac{\partial \mu_i}{\partial c} = -RT\bar{V}_i \left(\frac{1}{M} + 2A_2 c + 3A_3 c^2 + \dots \right) = -V_i^\circ \left(\frac{\partial \pi}{\partial c} \right)$$

$$\text{Then } \frac{-1}{kT\bar{V}_i} \left(\frac{\partial \mu_i}{\partial c} \right) = N \left(\frac{1}{M} + 2A_2 c + 3A_3 c^2 + \dots \right) = \frac{-N(\partial \pi)}{RT}$$

$$\frac{I_s}{\sqrt{I_i}} = \frac{4\pi^2 n_0^2 (\delta n/c)^2}{\lambda_0^4 r^2} \frac{c}{N \left(\frac{1}{M} + 2A_2 c + 3A_3 c^2 + \dots \right)}$$

OPTICAL STUFF THERMODYNAMIC STUFF

$$\text{Or... } \frac{I_s}{\sqrt{I_i}} = \frac{4\pi^2 n_0^2 (\delta n/c)^2}{N \lambda_0^4 r^2} \frac{c}{\frac{1}{RT} (\delta n/c)}$$

So... The LS experiment gives you everything De osmometer does (for monodisperse samples anyway).

This is often rearranged into new terms.

Define: $R = \frac{r^2}{\sqrt{I_i}} \frac{I_s}{c}$ = Rayleigh factor {not "ratio" as is}
 often said

UNITS: cm⁻¹

$$K = \frac{4\pi^2 n_0^2 (\delta n/c)^2}{\lambda_0^4 N} = \text{"optical constant"}$$

Then we have $R = \frac{K c}{\frac{1}{RT} (\delta n/c)}$

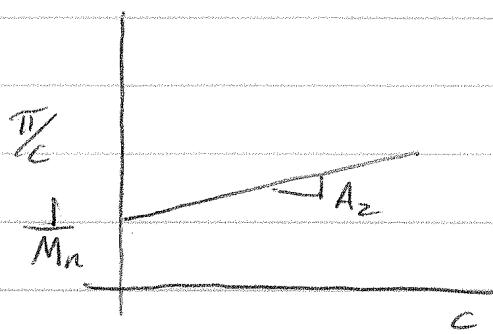
or... $\frac{K c}{R} = \frac{1}{RT} (\delta n/c) = \frac{1}{RT} \left(RT \left(\frac{1}{M} + 2A_2 c + \dots \right) \right)$

easily memorized forms

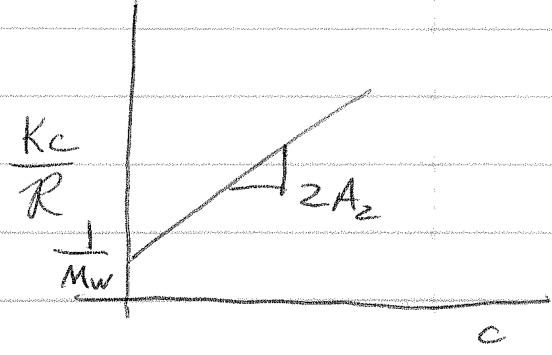
or $\frac{K c}{R} = \frac{1}{M} + 2A_2 c + \dots$

Graphically, with comparison to osmometer & limitations & characteristics.

OSMOMETER



L.S.



Limits

$$10^4 \leq M_z \leq 10^6$$

↑
due to
membrane

↑
due to
sensitivity

$$500 \leq M_w \leq 10^8$$

↑
due to
sensitivity

↑
due to need to
go to low angles
(see below)
depends on how
good you are!

CHARACTERISTICS

Responds slowly to energy required to make large, macroscopic pressure changes, distance scale \approx cm.

CHARACTERISTICS

Responds instantly to small energies required to thermally stimulate small pressure changes over microscopic distances (~~length~~ ($\approx \lambda/n$, see below)).

Problem: - Leaks

- not easily automated/coupled to other techniques

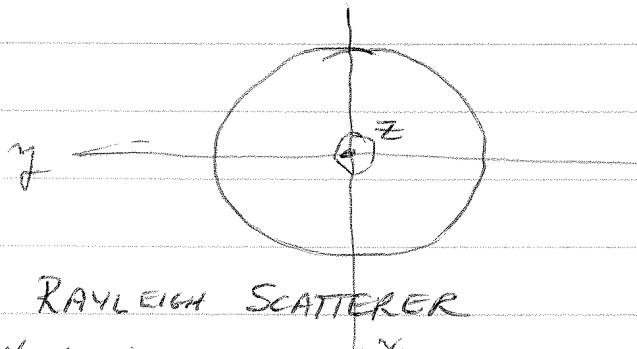
Problem: dust, finite size.

EXTENSION TO LARGE PARTICLES

The development so far is due to some of the greatest minds in science - Einstein, Smoluchowski, Debye, Rayleigh - and much of the development proceeds belief in large molecules.

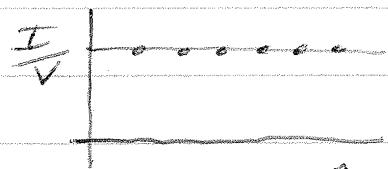
But surely the observation concerning the unusual behavior - distorted donut - of large particles must have been made.

Looking down on a donut, z polarized incident light

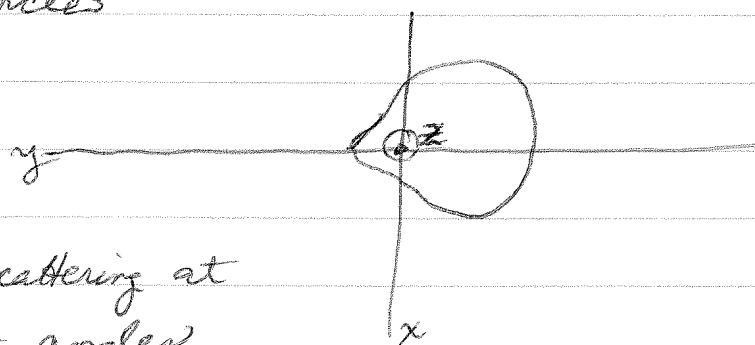


Small particle: RAYLEIGH SCATTERER

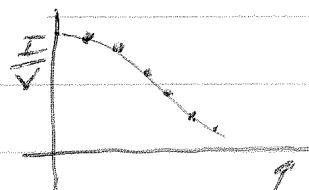
uniform scattering in
all directions in the horizontal plane



Large particles



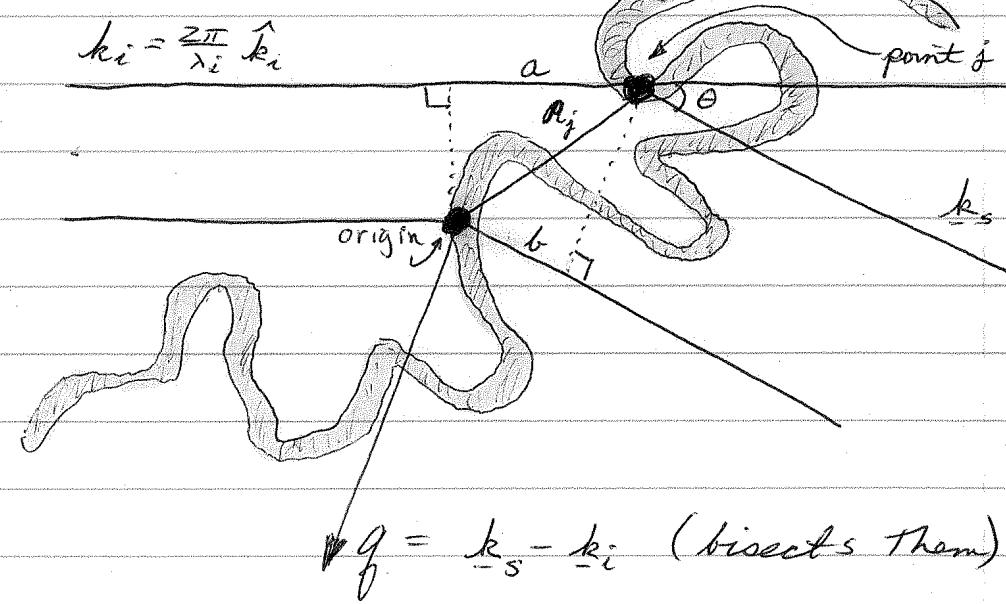
reduced scattering at
larger angles.



6

Extension to Large Particles

$$\underline{k}_i = \frac{2\pi}{\lambda_i} \hat{\underline{k}}_i$$



$$\underline{k}_s = \frac{2\pi}{\lambda_s} \hat{\underline{k}}_s$$

$\nabla g = \underline{k}_s - \underline{k}_i$ (bisects them)

Note: all λ 's are λ_0/n here, where n = solution refractive index
it is assumed that $\lambda_i = \lambda_s$

extra distance of ray through point j compared to origin:

$$\begin{aligned}
 \text{phase shift} &= \frac{\nabla g \cdot \underline{R}_j}{2\pi} = a - b \\
 &= (\underline{R}_j \cdot \underline{k}_i) - (\underline{R}_j \cdot \underline{k}_s) \\
 &= -g \cdot \underline{R}_j
 \end{aligned}$$

LS-35

Before we start adding fields & squaring, let's look at $|\underline{q}|$

It turns out that the magnitude of \underline{q} is useful because we can often assume that molecules re-orient randomly & rapidly w/r/s. to \underline{q} .

$$\begin{aligned} |\underline{q}| &= |\underline{q}_\perp| = \sqrt{(\underline{k}_i - \underline{k}_s) \cdot (\underline{k}_i - \underline{k}_s)} \\ &= \sqrt{k_i^2 - 2k_i \cdot k_s \cos\theta + k_s^2} \\ &= \sqrt{\left(\frac{2\pi}{\lambda}\right)^2 (2 - 2 \cos\theta)} \\ &= \frac{2\pi}{\lambda} \sqrt{2(1 - \cos\theta)} \end{aligned}$$

But fig. tells us that $1 - \cos\theta = 2\sin^2\frac{\theta}{2}$

$$\text{So... } q = \frac{2\pi}{\lambda} \sqrt{2(2\sin^2\frac{\theta}{2})} = \frac{4\pi}{\lambda} \sin \frac{\theta}{2}$$

Recall $\lambda = \lambda_0/n$

$$\text{So... } q = \frac{4\pi n}{\lambda_0} \sin \frac{\theta}{2}$$

Other symbols for q : K , S , h , a

Now we must add & square the fields E_j from all the N scatterers in the molecule

$$E_j = E_0 e^{i(\omega t + \phi)} = E_0 e^{i(\omega t - q \cdot R_j)}$$

$$E_s = \sum_{j=1}^N E_j = E_0 e^{i\omega t} \sum_{j=0}^N e^{-iq \cdot R_j}$$

$$I_s(\theta) = |E_s|^2 \sum_{i,j}^N e^{iq \cdot (r_i - r_j)}$$

call this R_{ij}

$q=0$ CASE Clearly, from this equation, at $q=0$ we get:

$$I_s(0) = |E_0|^2 \sum_{i=1}^N \sum_{j \neq i}^N 1 = N^2 |E_0|^2$$

Note how different this is from our finite angle gas expression where we added and then squared the ~~phases~~ fields of many particles, scattered to finite ~~and different~~ angles. Because those particles can change their ϕ at finite q , we can "wipe out" the off-diagonal terms. Here, no matter how the scattering segments move, at $q=0$ they stay in phase.

I think this bears a formal resemblance to the "shot noise" term in DLS

$q = \text{finite case}$ Anticipation: we won't get exactly zero as in the independent particles case because particles ~~are~~ are correlated now due to connectedness.

Define $P(\theta) = \frac{I_s(\theta)}{I(0)}$ So $P(\theta)$ is always ≤ 1 since $I(0)$ is the inphase intensity.

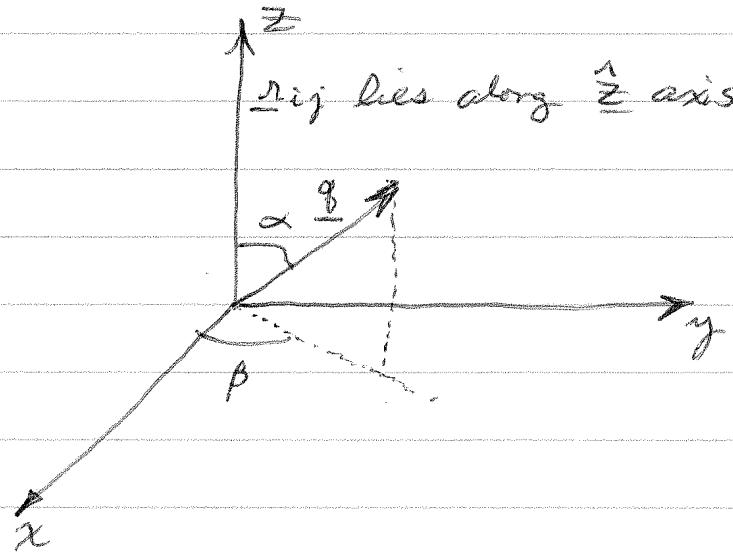
A note:
even at this point the inverse nature of LS is evident:
for $q=0$ see nothing
 $high q \Rightarrow high magnification$

$$\Rightarrow P(\theta) = \frac{1}{N^2} \sum_{i,j=1}^N e^{iq \cdot (r_i - r_j)}$$

Orientational Averaging

The foregoing expression gives $P(\theta)$ for one orientation of \underline{r}_{ij} with respect to \underline{q} . But in normal solutions, \underline{r}_{ij} can reorient freely. Typically, many reorientations will occur in the time required to complete an intensity measurement, so we should perform an average of \underline{r}_{ij} w/r respect to some fixed \underline{q} .

Equivalently, we could let \underline{r}_{ij} "stand still" and let \underline{q} assume all orientations:



This kind of problem occurs often in physics.

We have to average out each term involving $g_{\alpha ij}$ over all the possible orientations between \underline{q} and r_{ij} .

Average of
one such term = $\int_{\beta=0}^{2\pi} \int_{\alpha=0}^{\pi} e^{-iq \cdot r_{ij}} \sin \alpha d\alpha d\beta$

 4π

This eqn. comes from $\iiint \sin \alpha d\alpha d\beta r^2 dr = \frac{4}{3}\pi R^3$ {
i.e. $dV = r^2 \sin \alpha d\alpha d\beta$ for spherical polar coordinates}
We do not vary r in this problem, since r_{ij} is fixed in magnitude & only varies by orientation.

$$\Rightarrow \frac{1}{4\pi} \int_0^{2\pi} d\beta \int_0^{\pi} e^{iq \cdot r_{ij} \cos \alpha} \sin \alpha d\alpha$$

$$\Rightarrow \frac{1}{2} \int_0^{\pi} e^{iq \cdot r_{ij} \cos \alpha} \sin \alpha d\alpha$$

$$\text{let } iq \cdot r_{ij} \cos \alpha = u$$

$$du = (-\sin \alpha)(iq \cdot r_{ij}) d\alpha$$

$$\text{So... } \sin \alpha d\alpha = \frac{du}{-iq \cdot r_{ij}}$$

$$-iq \cdot r_{ij}$$

$$\Rightarrow \frac{1}{2} \cdot \frac{1}{(-iq \cdot r_{ij})} \cdot \int e^u du$$

$$\Rightarrow \frac{-1}{2iq \cdot r_{ij}} \underbrace{e^{iq \cdot r_{ij} \cos \alpha} \Big|_0^\pi}_{e^{-iq \cdot r_{ij}} - e^{+iq \cdot r_{ij}}}$$

Now recall from complex variables that $\sin x = \frac{e^{ix} - e^{-ix}}{2i}$

$$\text{Thus, } e^{-iqr_{ij}} - e^{+iqr_{ij}} = -2i \sin q r_{ij}$$

So... "average of one such term in $P(\theta)$ " = $\frac{\sin q r_{ij}}{q r_{ij}}$

So... at last!

$$P(\theta) = \frac{1}{N^2} \sum_{i,j=1}^N \frac{\sin q r_{ij}}{q r_{ij}}$$

$$\text{where } q = \frac{4\pi n}{\lambda_0} \sin \frac{\theta}{2}$$

$$r_{ij} = |\underline{r}_i - \underline{r}_j|$$

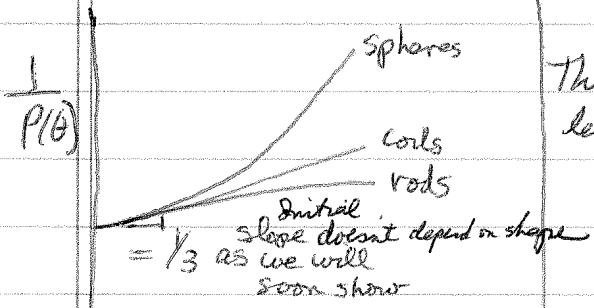
$N = \# \text{ scattering elements}$

It was assumed that all r_{ij} can orient w/r to q randomly

Now there are 2 options, depending on a priori shape information and monodispersity info and overall size.

- Particle shape is known, or sample known monodisperse. Then it is possible to develop explicit equations for $P(\theta)$.

Examples



Thin rod length L $P(\theta) = \frac{2}{\pi} \int_0^x \frac{\sin u}{u} du - \left(\frac{\sin(x/2)}{(x/2)} \right)^2$

$$x = gL$$

$$R_g^2 = \frac{L^2}{12}$$

Some if you know
That the sample is
monodisperse, you
can generate $P(\theta)$
curves & see if they match
experiment!

Uniform sphere radius R $P(\theta) = \left(\frac{3(\sin x - x \cos x)}{x^3} \right)^2$

$$x = gR$$

$$R_g^2 = \frac{3}{5} R^2$$

Gaussian coil* $P(\theta) = \frac{2}{x^2} (e^{-x} - 1 + x)$
mean squared end to end length, $\langle r^2 \rangle$ $x = g^2 \langle r^2 \rangle = g^2 R_g^2$

$$\therefore R_g^2 = \langle r^2 \rangle$$

* This "Dilys equation" seems to work OK also for expanded coils that are not too large.

- Particle shape is not known, or sample polydisperse. Then it is still possible to obtain a radius of gyration, provided that measurements can be made at sufficiently low gR_g .

- Shape not known, continued.

To get R_g , regardless of shape, we obtain the limiting form of $P(\theta)$ at low g .

$$\text{We had } P(\theta) = \frac{1}{N^2} \sum_{i,j} \frac{\sin g r_{ij}}{g r_{ij}}$$

$$\text{but } \sin x \approx x - \frac{x^3}{3!} + \frac{x^5}{5!} \text{ at low } x, \text{ i.e. low } g r_{ij}$$

$$\text{So } \frac{\sin x}{x} \approx 1 - \frac{x^2}{6}$$

$$P(\theta) = \frac{1}{N^2} \sum_{i,j} \left(1 - \frac{g^2}{6} r_{ij}^2 \right)$$

$$\text{But } \sum_{i,j} 1 = N^2$$

$$\text{and } \sum_{i,j} r_{ij}^2 = 2N^2 R_g^2 \quad \left\{ \begin{array}{l} \text{See Tanford Sec.} \\ q_j, \text{ eqn 9-34} \\ p.167 \end{array} \right.$$

Inverse Plot

$$\text{so } P(\theta) = 1 - \frac{g^2 R_g^2}{3} + \dots$$

$$\text{or... } \frac{1}{P(\theta)} \approx 1 + \frac{g^2 R_g^2}{3} \text{ which explains the initial slope in figure on previous page.}$$

$$\text{If polydisperse, get Z average of } R_g^2 : R_g^2 = \frac{\sum n_i M_i^2 R_g^2}{\sum n_i M_i^2}$$

Guinier Plot

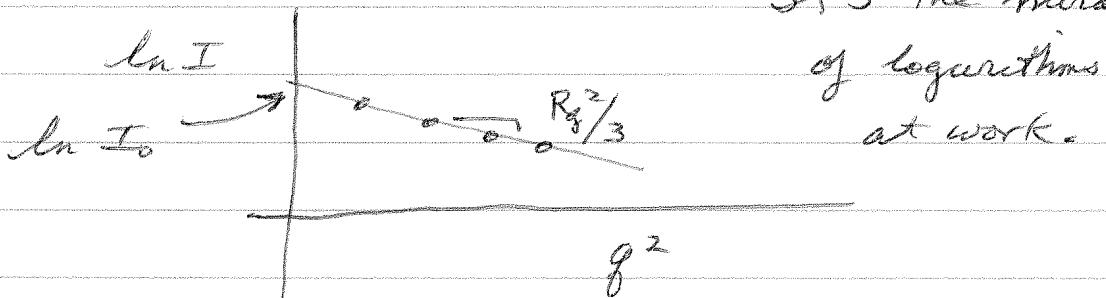
This limiting form has an important alternative form.

$$P(q) \approx 1 - q^2 R_g^2 \approx e^{-q^2 R_g^2 / 3}$$

$$\text{Thus } \ln P(q) = \ln \left(\frac{I}{I_0} \right) = \ln I - \ln I_0 = -\frac{q^2 R_g^2}{3}$$

The important advantage of this form is that I_0 need not be estimated to get R_g ! Just plot $\ln I$ vs. q^2 .

It's the miracle



Limitations

In order to use either Debye or Guinier plots, $I(q)$ must actually be decreasing. This will be a sensitive function of radius because the decreasing term is the square of $q R_g$. As a practical guide, R_g should be bigger than $\approx \lambda/20$. If not, you have to be very careful ... or consider another technique (dynamic LS or

Laser	$\lambda/\text{\AA}$	λ in toluene (\AA)	$\lambda/20$	X-ray or neutron scattering?
Ar ⁺	4880	32.50	163	
Ar ⁺	5145	34.30	17.2	
HeNe	6328	42.20	21.0	

PUTTING IT ALL TOGETHER

LS-47

THE WRONG WAY (However, you can find it
in respected books - e.g. Kerner)
also van Hove

"Pasting" The finite size results onto our previous thermodynamic results, which here would be valid only at $g = 0$.

Before we did finite size, we had: $\frac{K_c}{R} = \frac{1}{M} + 2A_2 c + \dots$

$$\text{recall that } R = \frac{r^2 I_s}{V I_i} \text{ i.e. } R \sim I_s$$

$$\text{So } \frac{K_c}{R} \sim \frac{1}{I_s} \sim \frac{1}{I_s(0) P(\theta)}$$

So we could "patch up" The small scatterer equation like this:
guess

$$\frac{K_c}{R} \stackrel{?}{=} \frac{1}{P(0)} \left(\frac{1}{M} + 2A_2 c + \dots \right)$$

$$\stackrel{?}{=} \frac{1}{\left(1 - \frac{g^2 R g^2}{3} \right)} \cdot \left(\frac{1}{M} + 2A_2 c + \dots \right)$$

$$\stackrel{?}{=} \left(\frac{1}{M} + 2A_2 c + \dots \right) \left(1 + \frac{g^2 R g^2}{3} + \dots \right)$$

It's wrong ... but it's close, and it even behaves OK in the $c=0$ or $g=0$ limits:

$$\lim_{g \rightarrow 0} \frac{K_c}{R} = \frac{1}{M} + 2A_2 c + \dots$$

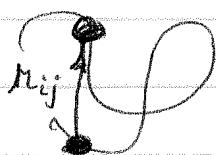
$$\lim_{c \rightarrow 0} \frac{K_c}{R} = \frac{1}{M} \left(1 + \frac{g^2 R g^2}{3} + \dots \right)$$

PUTTING IT ALL TOGETHER

THE RIGHT WAY (Zimm, J. Chem. Phys. 16 (12), 1093-1116, 1948)
(two papers)

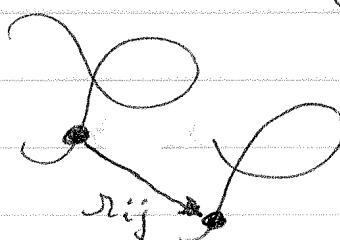
What he did: $I_s = \sum e^{i\theta \cdot \underline{r}_{ij}}$
all possible
segments, whether on
same polymer or not

So, this way you can include these interferences
(intramolecular)



as well as these ones

(intermolecular)



So... That's pretty hard. Assumption:

Assumptions:

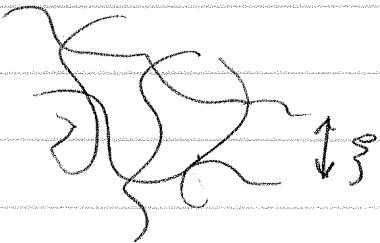
Result:

$$\frac{K_c}{R} \approx \frac{1}{M P(\theta)} + 2A_2 c$$

The $P(\theta)$ term
applies only to
the intercept, not
to the virial term.

A MODERN CHOICE

forget molecular parameters. Leave it as a thermodynamic expression, modified by a "correlation length" representing the persistence of phase among all scatterers, whatever chain they may be on.



$$\frac{K_c}{R} = \frac{1}{RT} \left(\frac{\partial \pi}{\partial c} \right)_{T, P} (1 + g^2 \xi^2)$$

$$\lim_{c \rightarrow 0} \xi^2 = \frac{Rg^2}{3}$$

$$\lim_{g \rightarrow 0} \frac{1}{RT} \left(\frac{\partial \pi}{\partial c} \right)_{T, P} = \frac{1}{M}$$